54th European Space Sciences Committee Plenary Meeting



Status of the Science Programme

Luigi Colangeli Head – Science Coordination Office Directorate of Science

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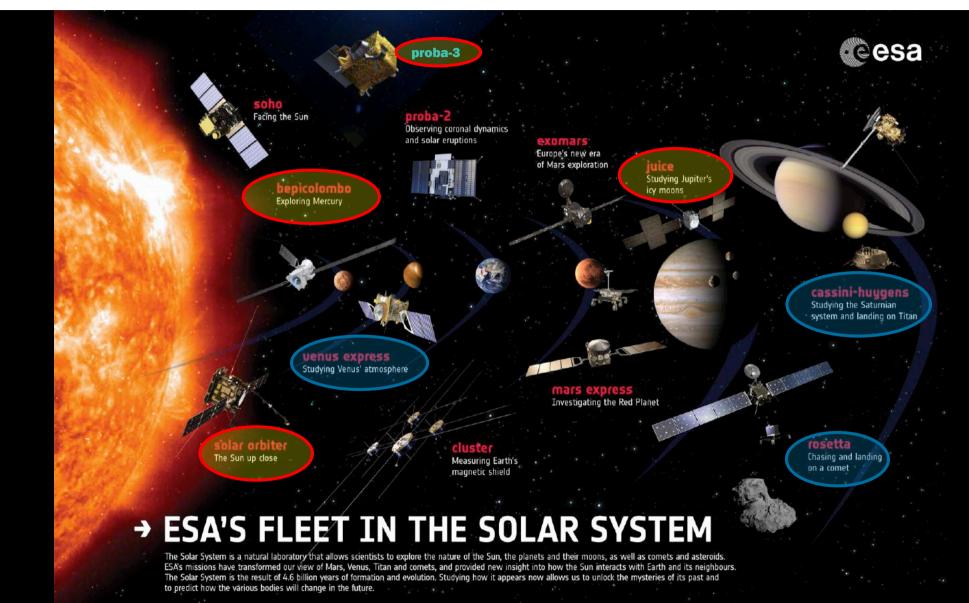
Overall status

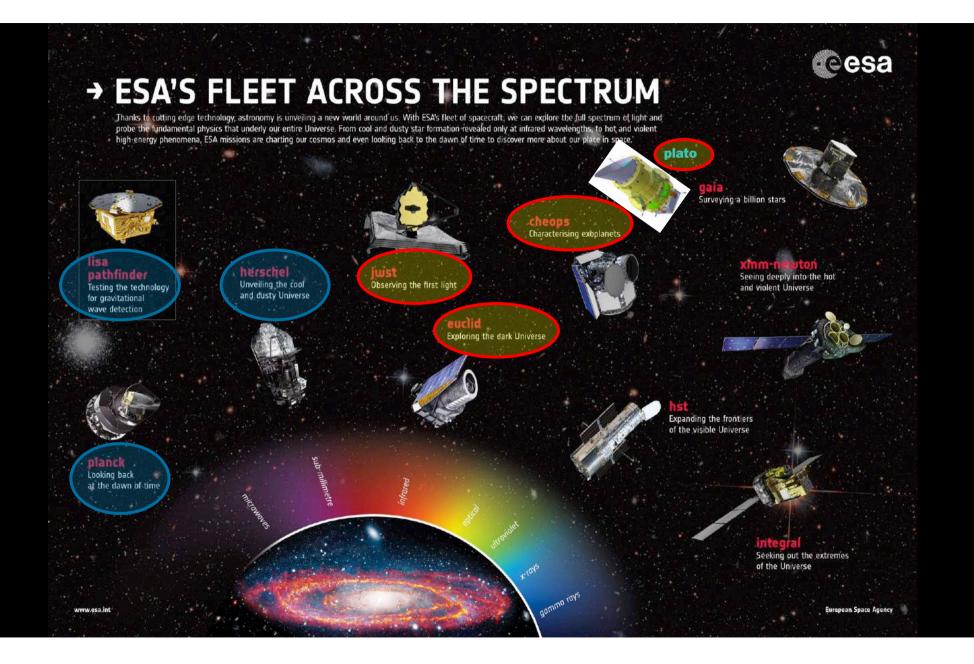
> M4 and M5 selection processes – status

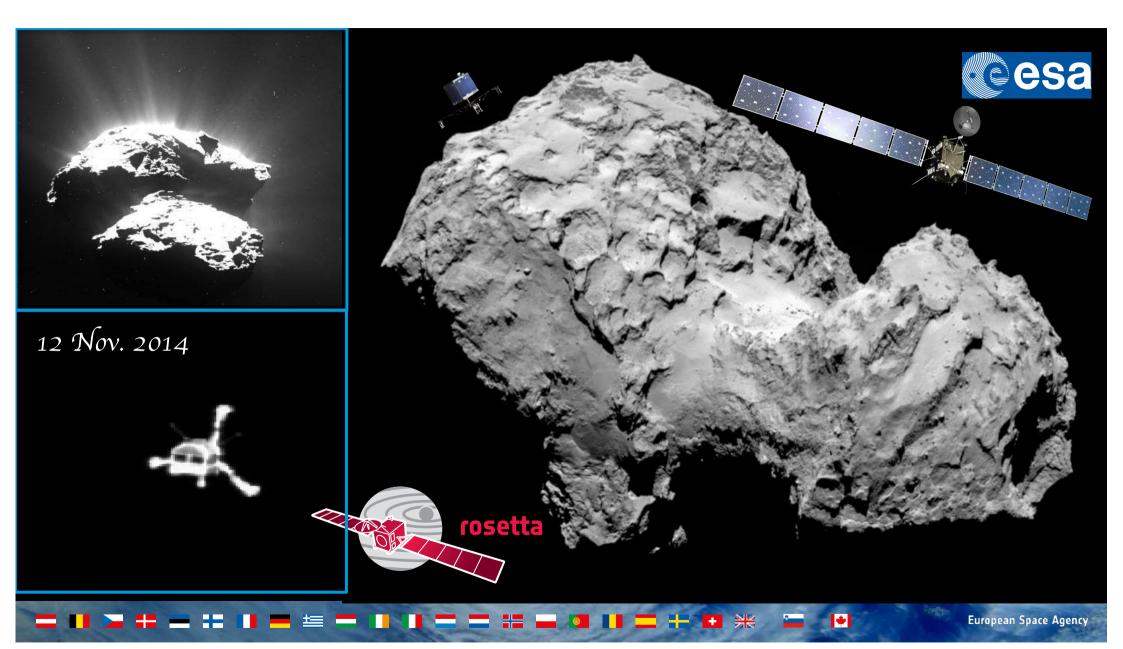
"New ideas"

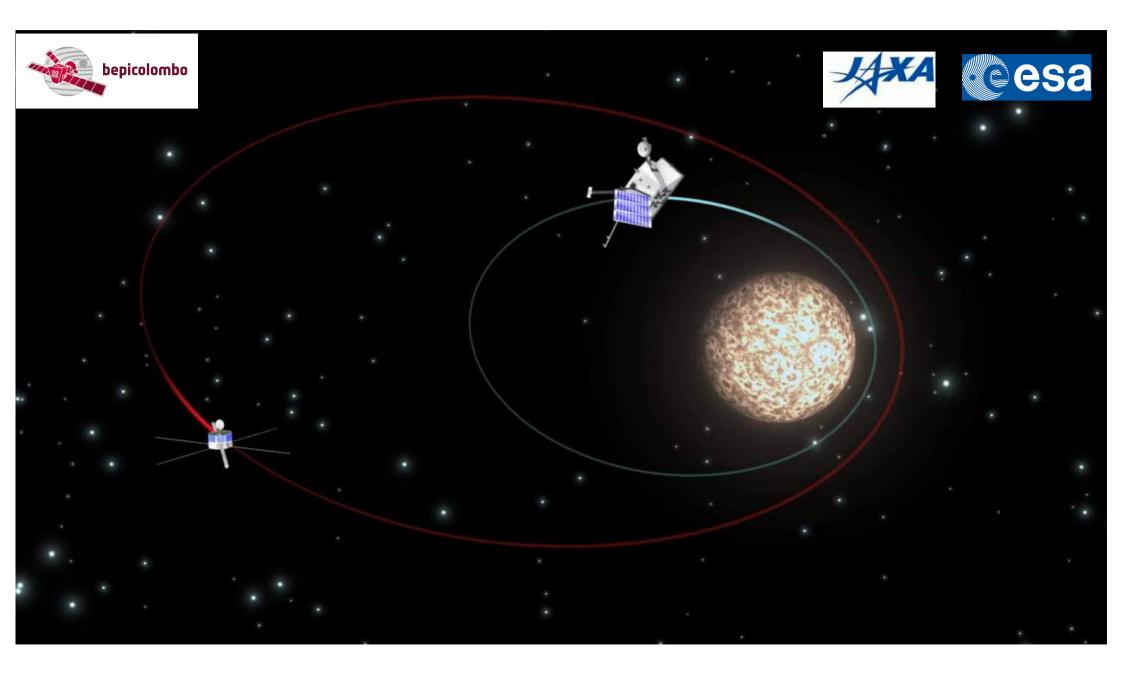
Other activities

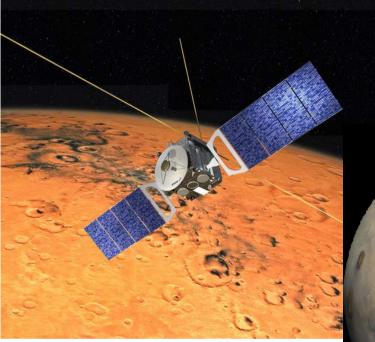




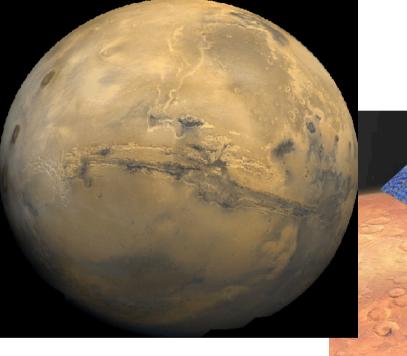










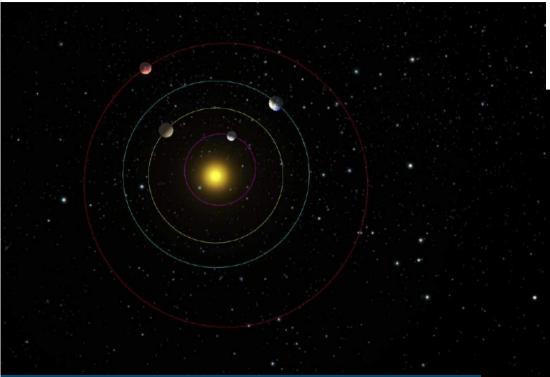




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Jupiter's Earth's Moon 10 3,476 Km 3,630 Km

Mercury 4,878 Km

Jupiter's

Europa

3138 Km





Ganymede

5,262 Km



Jupiter's Callisto 4,800 Km

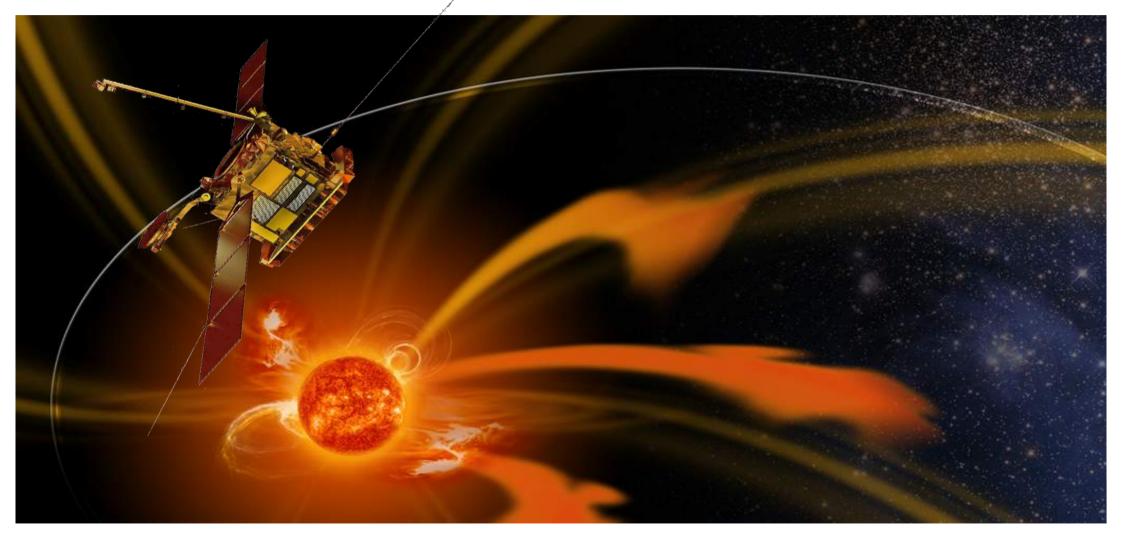
Saturn's Titan 5,150 Km

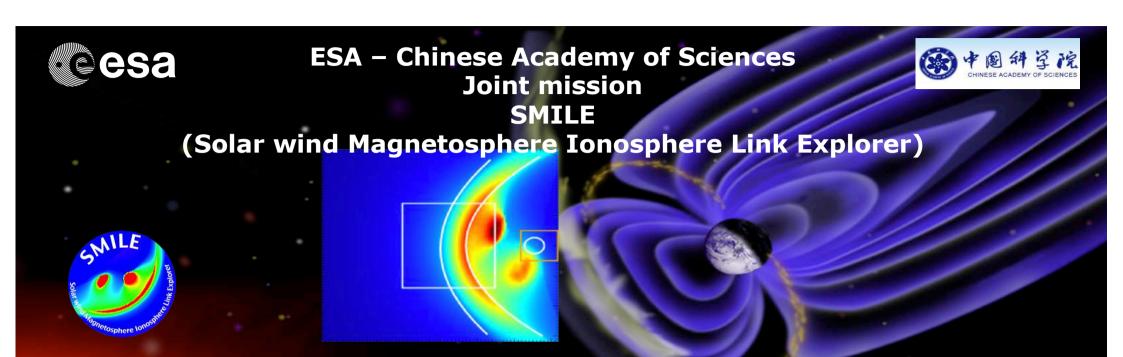


Solar Orbiter: Exploring the Sun-Heliosphere Connection









Co-Pis: G. Branduardi-Raymont and C. Wang

Smile will investigate the interaction between Earth's protective shield – the magnetosphere – and the supersonic solar wind

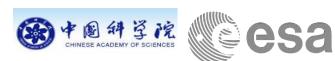
Goal: understanding the physical processes taking place during the continuous interaction between the solar wind and the magnetosphere

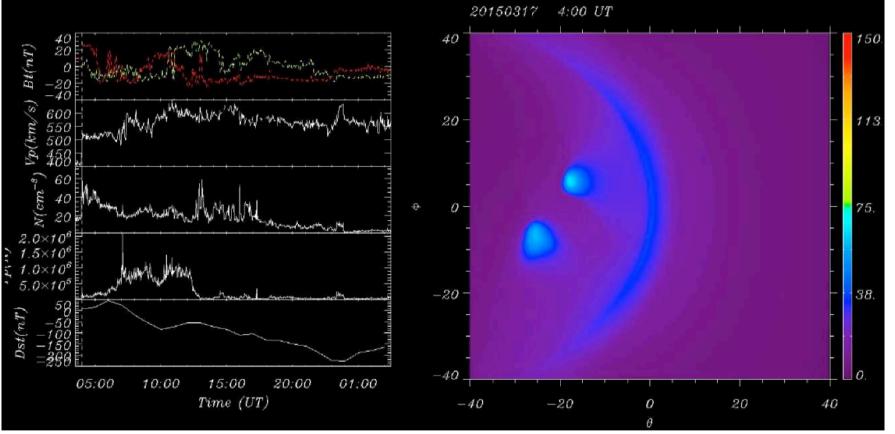
Aurora: NASA Polar



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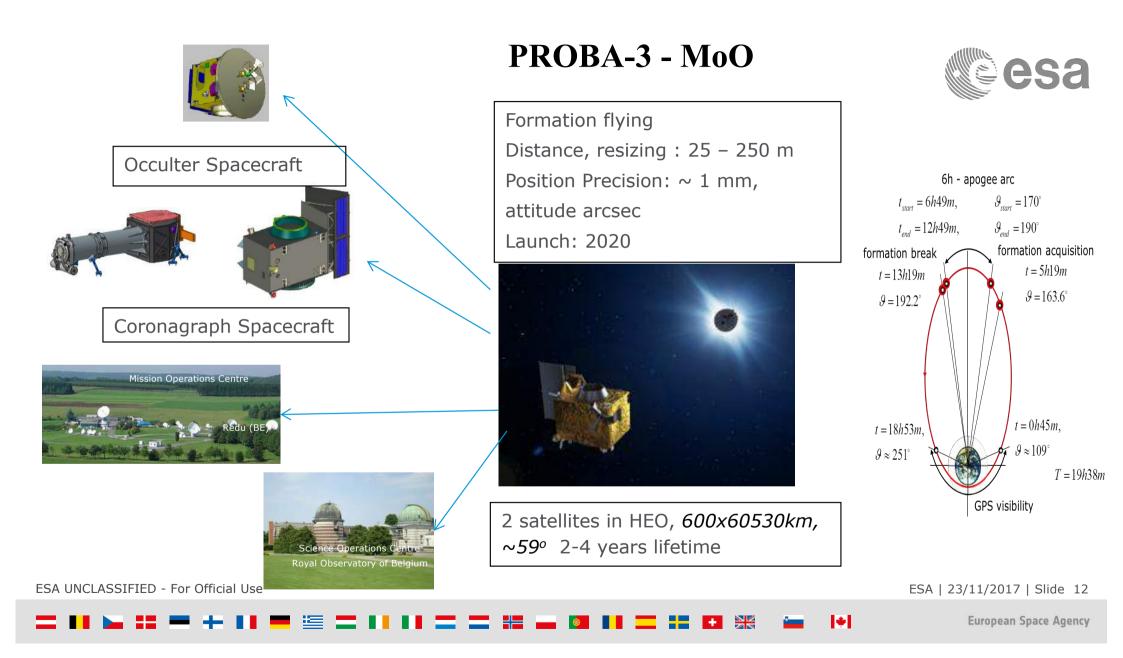
Simulated SMILE X-ray movie: 17th March 2015 solar storm, Earth impact

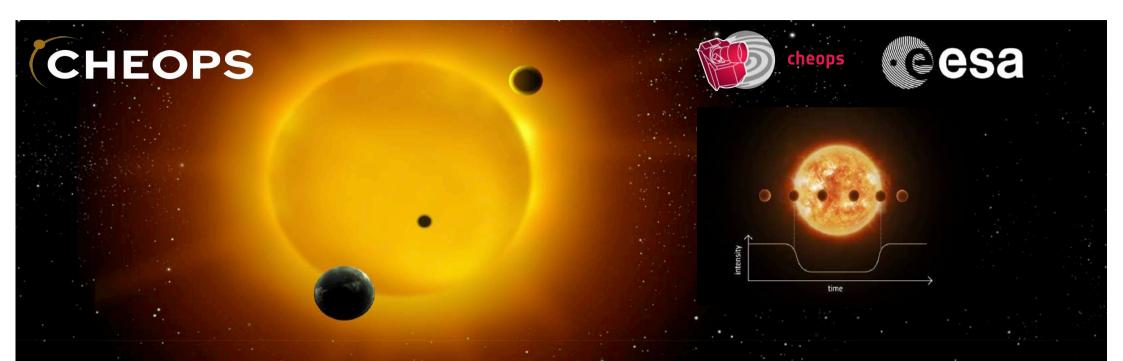




T. Sun, NSSC, CAS ESA | 23/11/2017 | Slide 11

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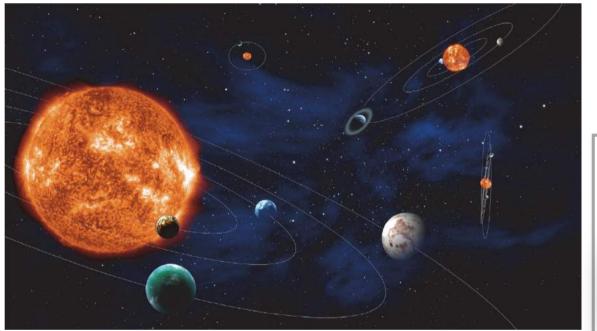




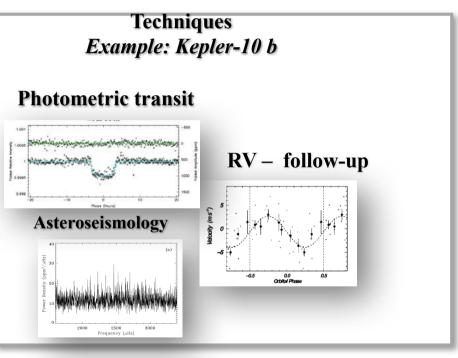
CHEOPS (CHaracterising ExOPlanet Satellite) – a mission dedicated to study exoplanet transits of local, bright stars already known to host exoplanets:

- Characterisation of transiting exoplanets with masses $< 30 M_{earth}$ through precision, wideband transit photometry
- Follow-up, pointed observations
 - Pointed observations
 - Bright host stars (V<12)
- Characterisation of super-Earths and Neptunes: precision $\underline{\text{masses} + \text{radii}} \rightarrow \text{measurement}$ of bulk density
- Identification of "golden targets" for spectroscopic characterisation
- Probing atmospheres of hot-Jupiters using phase curve measurements

PLATO mission – M3 - Adopted



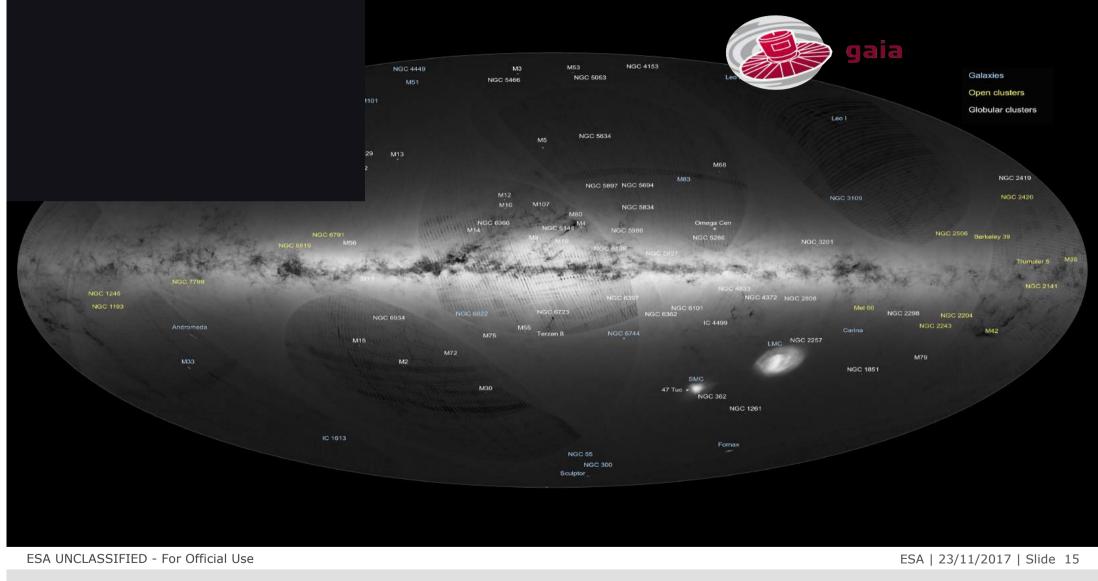




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European Space Agency

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NASA flagship mission

partnership between NASA, ESA and the Canadian Space Agency

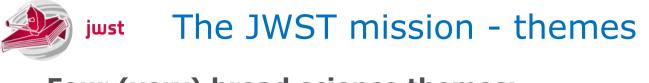
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- General purpose near and mid-infrared observatory
- Largest astronomical telescope ever flown
- Observing objects ranging from planets and bodies of our Solar System to some of the most distant galaxies



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Four (very) broad science themes:

The end of the dark ages: first light and reionisation.

The assembly of galaxies.

The birth of stars and proto-planetary systems.

Planetary systems (including our Solar System) and the origin of life.

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See Gardner et al., 2006, Space Science Reviews, 123, 485



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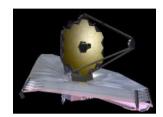


JWST is a partnership between NASA, ESA and CSA

ESA (and CSA) have been present since the very early phases of the mission

In return for this contribution, ESA shall obtain a portion of the observing time on JWST that will be no less than 15% of the observing time on average over the lifetime of the mission.











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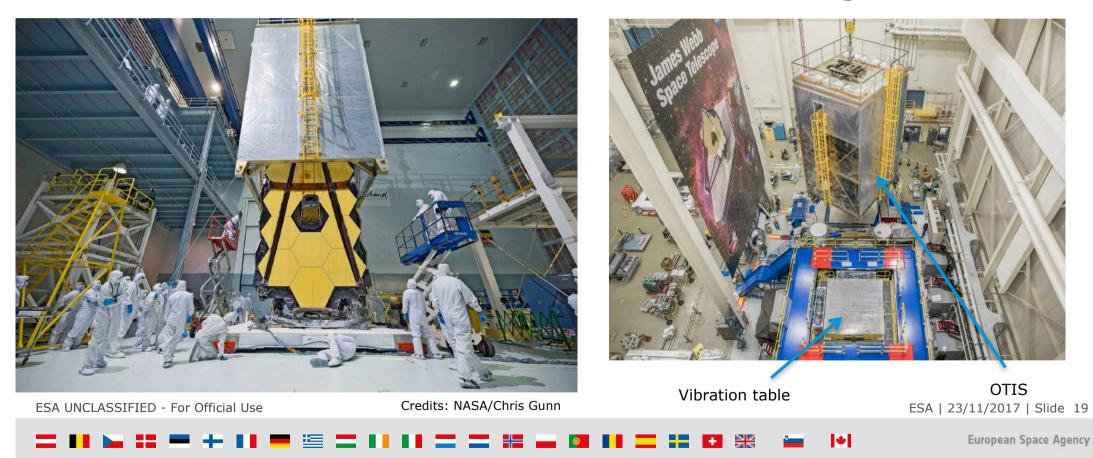
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jwst JWST hardware status - OTIS



Making sure the telescope and the instruments can survive the harsh conditions of a rocket launch: acoustic and vibration testing.



X-ray Astronomy Recovery Mission (XARM) - MoO

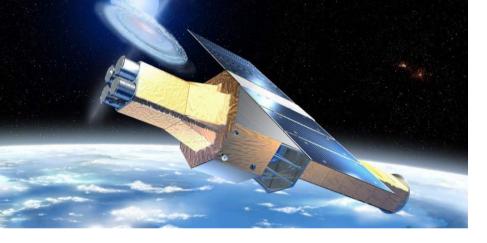


- > Hitomi (pioneering X-ray micro-calorimeter science) was lost 6 weeks after launch.
- > 2 Nature papers on the first ~ 1 week-long observation of the Perseus Cluster.
- JAXA Recovery mission
 - Micro-calorimeter (*Resolve*), ≤7 eV energy resolution.
 3'x3' field-of-view
 - Large-field (~40'x40') CCD detectors (*Xtend*),
 ≤170 eV resolution @6 keV
 - Soft X-ray telescope, ~1.3' Half Energy Width
- ESA MoO + CH and NL national contributions
- Return to European scientists:
 - Membership of the Science Team as Participating Scientists (access to PV data) Call will be issued shortly
 - 8% (TBC) of Open Time for European proposals.

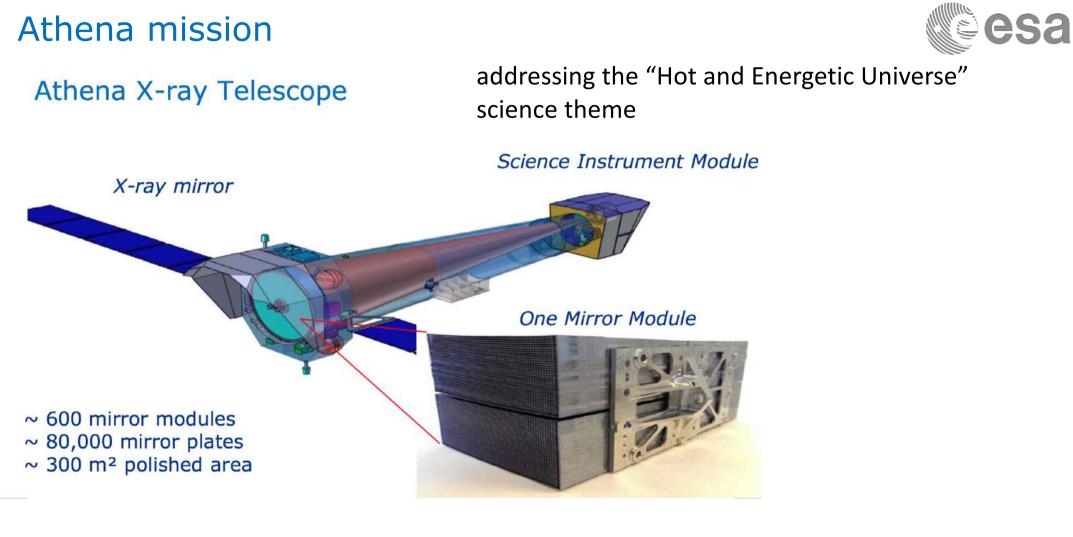
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Differential Acceleration noise (as published)

→ LISA PATHFINDER EXCEEDS EXPECTATIONS

being investigated



Cesa

10-12 Residual relative acceleration of the test masses [ms⁻² / VHz] 10-13 10-14 Frequency [Hz] 10-15 0.0001 0.001 0.010.11 **Centrifugal force** Gas damping Sensing noise The rotation of the spacecraft Inside their housings, the test The sensing noise of the optical required to keep the solar array masses collide with some of the metrology system used to monitor few gas molecules still present. pointed at the Sun and the antenna the position and orientation of the pointed towards Earth, coupled with This noise term becomes smaller test masses, at a level of 35 fm / √Hz, the noise of the startrackers produces with time, as more gas molecules has already surpassed the level a noisy centrifugal force on the test are vented to space. of precision required by a future masses. This noise term has been gravitational-wave observatory by subtracted, and the source of the a factor of more than 100. residual noise after subtraction is still

LISA mission – L3 - Selected

- LISA is a mission to detect and observe gravitational waves
 - Gravitational wave astronomy, Fundamental physics, cosmology
- Laser interferometry to register changes in distance between test-masses moving on geodesics
- LISA Pathfinder successful technology demonstrator
 - Demonstrated ability to keep test-masses on geodesics with disturbances ${\sim} \text{fN}$
- System and payload CDF's completed, Mission Definition Review starting, industrial phase A April 2018, launch planned for 2034







The elements of the Science Programme



The building blocks of the programme include:

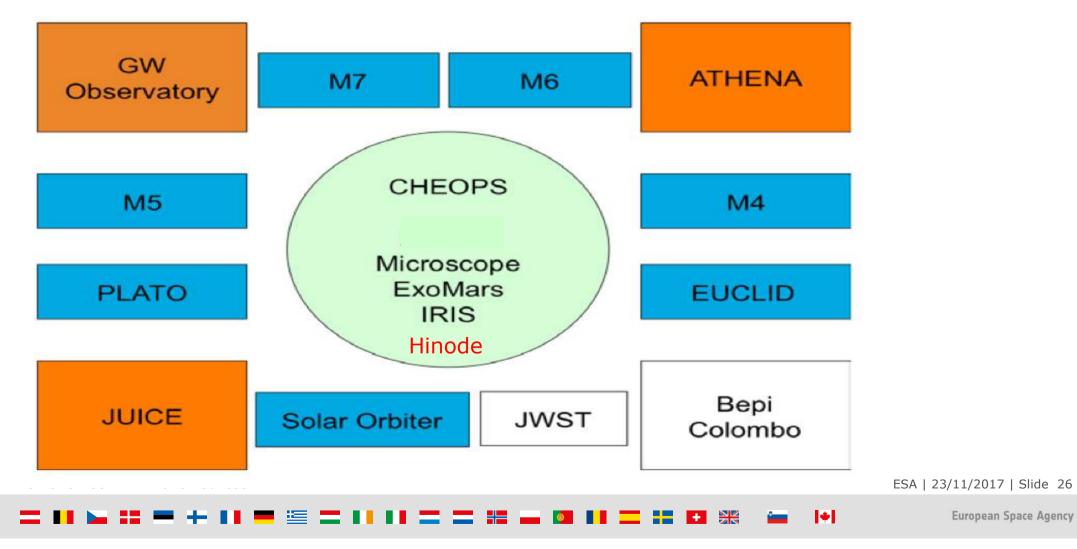
- **a.** L-missions, large European led flagship missions with a cost to ESA of around 2 annual budgets, one every 7-8 years.
- **b. M-missions**, provide the programme with flexibility. ESA led or implemented through international collaboration. Cost to ESA of around one annual budget, one every 3-4 years.
- **c. S-missions**, new concept allowing national agencies to play a leading role in missions, 0.1 annual budgets, one every 4 years, potentially.
- **d. O-missions**, which are "missions of opportunity", led by other agencies, small contributions.

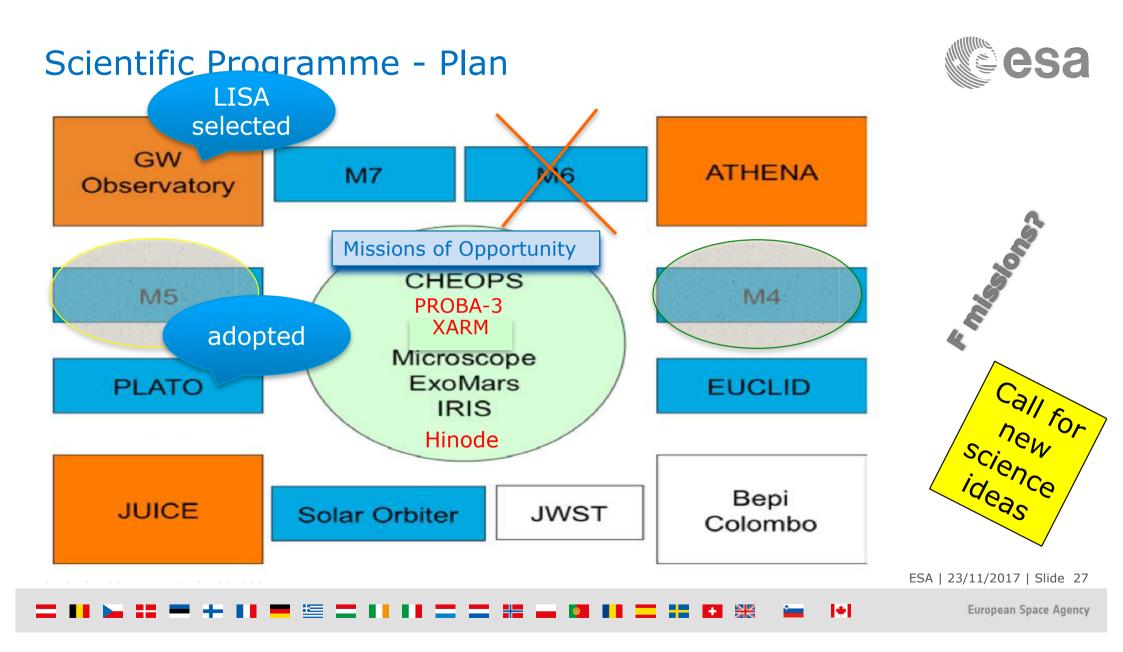




Scientific Programme - Plan







Recent SPC decisions (June & November 2017)



- Financial Situation and Projections of the Scientific Programme
- Adoption of the PLATO mission
- Selection of the L3 mission => LISA
- > Participation in the JAXA X-ray Astronomy Recovery mission (XARM) as MoO
- Participation in the PROBA-3 mission as MoO
- "Rules of the game" for Missions of Opportunity
- Extension of missions in operation for 2019-2020
- Resolution on the Chairship of Council's subordinate bodies

Science Programme Committee:

Chair: Prof. Stefano Vitale (IT)

Vice-Chair: Mr. Pauli Stigell (FI)



Mission "indicative" extension 2019-2020



- INTEGRAL from 1 January 2019 to 31 December 2019
- Gaia from 25 July 2019 to 31 December 2020
- Hinode*, HST*, IRIS*, Mars Express, SOHO and XMM-Newton from 1 January 2019 to 31 December 2020

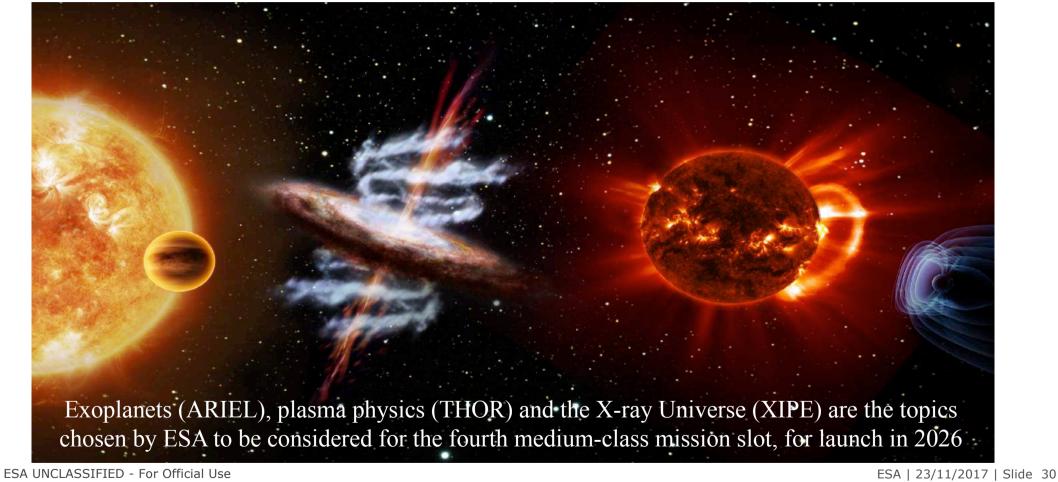
subject to a mid-term review and subsequent SPC confirmation in 2018

Cluster will be considered in February 2018, as well as INTEGRAL for 2020

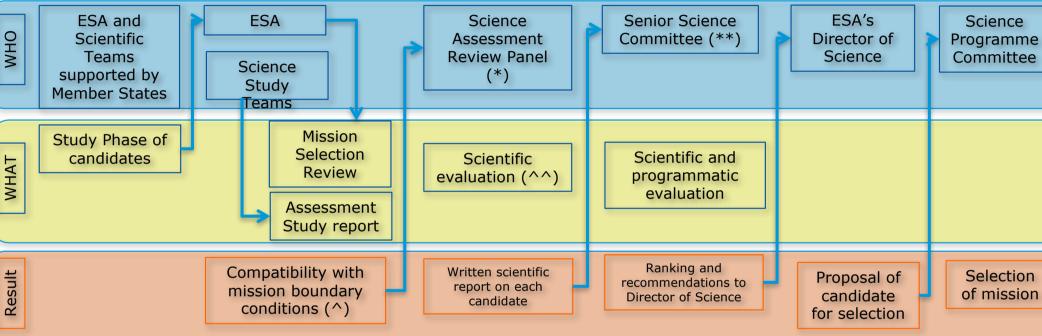


M4 selection





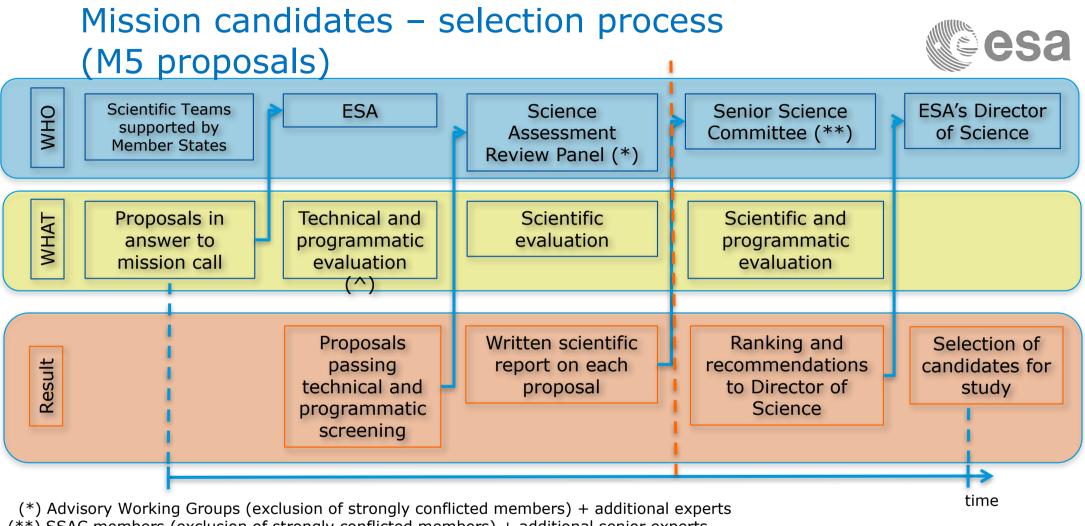
Mission – selection process (M4 candidates)



- (*) Advisory Working Groups (exclusion of strongly conflicted members) + additional experts
- (**) SSAC members (exclusion of strongly conflicted members) + experts
- (^) Including financial envelope, TRL of mission elements and readiness of Funding Agencies to fund mission elements proposed not to be under ESA's responsibility
- (^^) Including demonstrated capability to obtain the scientific objectives declared at the time of candidate selection

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time



(**) SSAC members (exclusion of strongly conflicted members) + additional senior experts

(^) Letters of Support from Funding Agencies

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New Science Ideas in ESA's Science Programme



- 26 proposals received by the deadline (14 September 2016)
- No a priori technical screening.
- Scientific assessment under the responsibility of the Advisory Structure, in two stages.
 - No prioritization, only identification of potentially interesting themes
- Post facto technical assessment
- Work on going on the three selected "themes"
- Results will be made public for the whole community



Selected themes (1/3) – Quantum Decoherence



- Strong interest in quantum physics (emphasis on boundaries of quantum decoherence, connection between gravitation and quantum physics)
- Science area with potentially high impact
- > Needs long, low-noise free fall -> ideally suited for LPF-like platform
- > P/L, science requirements, mission definition still immature
- Workshop (Trento, 6-7 June, 2017) with proposing teams and independent experts to better understand requirements, maturity, etc., to be followed by CDF study to assess mission maturity and identify areas for enabling technology developments
- Interaction with the community on going to prepare the CDF study target: early 2018



Selected themes (2/3) – Planetary science vs. platform size



- Strong interest in "focused" planetary missions based on small platforms
- Could enable significant additional opportunities for planetary science on rocky planets, small bodies
- Ideally suited for potential partnerships
- Workshop on "Planetary science missions vs. platform size" held on 6-7 September
- Interaction with the community => CDF started in November 2017



Selected themes – (3/3) High accuracy astrometry in the near IR



- Gaia-like mission in the near IR: beat reddening, access bulge, Galactic centre, hidden regions -> global view of the Milky Way
- Interaction with relevant community completed
- CDF study completed

